# Markscheme 

May 2018

Physics

Higher level

Paper 3

30 pages

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## Section A



| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 2. | a |  | combines the two equations to obtain result «for example $\frac{1}{I}=K^{2}(C+x)^{2}=\frac{4 \pi}{P}(C+x)^{2}$ » <br> $O R$ reverse engineered solution - substitute $K=2 \sqrt{\frac{\pi}{P}}$ into $\frac{1}{I}=K^{2}(C+x)^{2}$ to get $I=\frac{P}{4 \pi(C+x)^{2}}$ | There are many ways to answer the question, look for a combination of two equations to obtain the third one | 1 |
| 2. | b | i | extrapolating line to cross $x$-axis / use of $x$-intercept <br> OR <br> Use C $=\frac{y \text {-intercept }}{\text { gradient }}$ <br> OR <br> use of gradient and one point, correctly substituted in one of the formulae $\checkmark$ accept answers between 3.0 and 4.5 «cm» | Award [1 max] for negative answers | 2 |

(continued...)
(Question 2 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 2. | b | ii | ALTERNATIVE 1 <br> Evidence of finding gradient using two points on the line at least 10 cm apart $\checkmark$ <br> Gradient found in range: 115-135 or 1.15-1.35 $\checkmark$ <br> Using $P=\frac{4 \pi}{K^{2}}$ to get value between $6.9 \times 10^{-4}$ and $9.5 \times 10^{-4} « W$ » and POT correct $\checkmark$ <br> Correct unit, W and answer to 1,2 or 3 significant figures <br> ALTERNATIVE 2 <br> Finds $I\left(\frac{1}{y^{2}}\right)$ from use of one point ( $x$ and $y$ ) on the line with $x>6 \mathrm{~cm}$ and C from (b)(i) to use in $I=\frac{P}{4 \pi(C+x)^{2}}$ or $\frac{1}{\sqrt{I}}=K x+K C$ <br> Correct re-arrangement to get $P$ between $6.9 \times 10^{-4}$ and $9.5 \times 10^{-4}$ «W» and POT correct <br> Correct unit, W and answer to 1,2 or 3 significant figures $\checkmark$ | For 3rd marking point if no unit given, assume answer is in $W$ <br> Award [3 max] for an answer between 6.9W and 9.5W (POT penalized in 3rd marking point) <br> Alternative 2 is worth [3 max] | 4 |

(Question 2 continued)

| Question |  | Answers | Notes | Total |
| :--- | :--- | :--- | :--- | :--- | :---: |
| 2. | c | this graph will be a curve / not be a straight line $\checkmark$ <br> more difficult to determine value of $K$ <br> OR <br> more difficult to determine value of $C$ <br> OR <br> suitable mathematical argument $\checkmark$ | OWTTE |  |

## Section B

## Option A - Relativity

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 3. | a | i | 1.25c $\checkmark$ |  | 1 |
| 3. | a | ii | ALTERNATIVE 1 $\begin{aligned} & u^{\prime}=\frac{(0.50+0.75)}{1+0.5 \times 0.75} c \\ & 0.91 c \checkmark \end{aligned}$ <br> ALTERNATIVE 2 $\begin{aligned} & u^{\prime}=\frac{-0.50-0.75}{1-(-0.5 \times 0.75)} c \\ & -0.91 c \checkmark \end{aligned}$ |  | 2 |
| 3. | b |  | nothing can travel faster than the speed of light (therefore (a)(ii) is the valid answer) | OWTTE | 1 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 4. | a | $0.60 c$ <br> OR $1.8 \times 10^{8}\left\langle\mathrm{~m} \mathrm{~s}^{-1} »\right.$ |  | 1 |
| 4. | b | ALTERNATIVE 1 <br> time interval in the Earth frame $=90 \times \gamma=112.5$ minutes $\checkmark$ <br> «in Earth frame it takes 112.5 minutes for ship to reach station» <br> so distance $=112.5 \times 60 \times 0.60 \mathrm{c}$ $1.2 \times 10^{12} \text { «m» }$ <br> ALTERNATIVE 2 <br> Distance travelled according in the spaceship frame $=90 \times 60 \times 0.6 c$ $=9.72 \times 10^{11} \text { «m» }$ <br> Distance in the Earth frame «= $9.72 \times 10^{11} \times 1.25 »=1.2 \times 10^{12}$ «m» |  | 3 |

(continued...)
(Question 4 continued)

| Question |  |  | Answers |  | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 4. | c |  | signal will take $<112.5 \times 0.60 \Rightarrow$ » 67.5 «minutes» to reach Earth «as it travels at $c$ » <br> OR <br> signal will take $\begin{aligned} & \frac{1.2 \times 10^{12}}{3 \times 10^{8}}=» 4000 \text { «s» } \checkmark ~\end{aligned}$ <br> total time $«=67.5+112.5 »=180$ minutes or 3.00 h or $3: 00 \mathrm{am} \checkmark$ |  |  | 2 |
| 4. | d | i | line from event $E$ to $A$, upward and to left with A on ct axis (approx correct) $\checkmark$ line from event $A$ to $B$, upward and to right with $B$ on ct' axis (approx correct) <br> both lines drawn with ruler at 45 (judge by eye) $\checkmark$ |  |  | 3 |

(continued...)
(Question 4 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 4. | d | ii | ALTERNATIVE 1 <br> «In spaceship frame» <br> Finds the ratio $\frac{O B}{O E}$ (or by similar triangles on $x$ or ct axes), value is approximately $4 \checkmark$ <br> hence time elapsed $\approx 4 \times 90 \mathrm{mins} \approx 6 \mathrm{~h}$ «so clock time is $\approx 6: 00 » \checkmark$ | Alternative 1: <br> Allow similar triangles using $x$-axis or ct-axis, such as distance 2 from diagrams below distance 1 | 2 |

(Question 4 continued)
Question

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5. | a |  | quantity that is the same/constant in all inertial frames $\checkmark$ |  | 1 |
| 5. | b | i | spacetime interval $=27^{2}-15^{2}=504$ «m² $\downarrow$ |  | 1 |
| 5. | b | ii | ALTERNATIVE 1 <br> Evidence of $x^{\prime}=0$ $t^{\prime} «=\frac{\sqrt{504}}{c} »=7.5 \times 10^{-8} « \mathrm{~s} » \downarrow$ <br> ALTERNATIVE 2 $\begin{aligned} & \gamma=1.2 \checkmark \\ & t^{\prime} \text { «= } \frac{9 \times 10^{-8}}{1.2} »=7.5 \times 10^{-8} \text { «s» } \end{aligned}$ |  | 2 |
| 5. | c |  | observer B measures the proper time and this is the shortest time measured OR time dilation occurs «for B 's journey» according to A <br> OR <br> observer B is stationary relative to the particle, observer A is not $\checkmark$ |  | 1 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 6. | a |  | $\gamma «=\frac{3350}{938} »=3.57$ |  | 1 |
| 6. | b | i | $\begin{aligned} & \text { energy of pion }=(3350 \times 2)-6200=500 « \mathrm{MeV} » \checkmark \\ & 500^{2}=p^{2} c^{2}+140^{2} \checkmark \\ & p=480 « \mathrm{MeV} \mathrm{c}{ }^{-1} » \end{aligned}$ |  | 3 |
| 6. | b | ii | path of pion constructed in direction around 4-5 o'clock by eye | eg: | 1 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 7. | a | i | boundary inside which events cannot be communicated to an outside observer OR <br> distance/surface at which escape velocity $=c \quad \checkmark$ | OWTTE | 1 |
| 7. | a | ii | mass of black hole $=7.2 \times 10^{36}$ «kg» $\checkmark$ $« \frac{2 G M}{c^{2}}=» 1 \times 10^{10} « \mathrm{~m} » \checkmark$ |  | 2 |
| 7. | b |  | wherever S-2 is in orbit, time observed is longer than 5.0 s <br> when closest to the star S-2 periodic time dilated more than when at greatest distance $\checkmark$ Justification using formula or time is more dilated in stronger gravitational fields $\checkmark$ |  | 2 max |

## Option B — Engineering physics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 8. | a | i | an object's resistance to change in rotational motion <br> OR <br> equivalent of mass in rotational equations | OWTTE | 1 |
| 8. | a | ii | $\Delta \mathrm{KE}+\Delta \text { rotational } \mathrm{KE}=\Delta \mathrm{GPE}$ <br> OR $\begin{aligned} & \frac{1}{2} m v^{2}+\frac{1}{2} I \frac{v^{2}}{r^{2}}=m g h \checkmark \\ & \frac{1}{2} \times 0.250 \times v^{2}+\frac{1}{2} \times 1.3 \times 10^{-4} \times \frac{v^{2}}{1.44 \times 10^{-4}}=0.250 \times 9.81 \times 0.36 \\ & v=1.2<\mathrm{m} \mathrm{~s}^{-1} » \end{aligned}$ |  | 3 |
| 8. | a | iii | $\omega «=\frac{1.2}{0.012} »=100<\mathrm{rad} \mathrm{s}^{-1} » \checkmark$ |  | 1 |
| 8. | b | i | force in direction of motion $\checkmark$ <br> so linear speed increases |  | 2 |
| 8. | b | ii | force gives rise to anticlockwise/opposing torque on wheel $\checkmark$ so angular speed decreases | OWTTE | 2 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 9. | a |  | ALTERNATIVE 1 $\begin{aligned} & \text { «Using } \frac{V_{1}}{T_{1}}=\frac{V_{2}}{T_{2}} \text { " } \\ & V_{2}=\frac{47.1 \times(273+19)}{(273-12)} \\ & V_{2}=52.7 \text { « } \mathrm{m}^{3} » \end{aligned}$ <br> ALTERNATIVE 2 $\begin{aligned} & \text { «Using } P V=n R T \text { » } \\ & V=\frac{243 \times 8.31 \times(273+19)}{11.2 \times 10^{3}} \\ & V=52.6 \text { «m}{ }^{3} » \end{aligned}$ |  | 2 |
| 9. | b |  | $\begin{aligned} & W «=P \Delta V »=11.2 \times 10^{3} \times(52.7-47.1) \\ & W=62.7 \times 10^{3} « \mathrm{~J} » \end{aligned}$ | Accept $66.1 \times 10^{3} \mathrm{~J}$ if 53 used Accept $61.6 \times 10^{3} \mathrm{~J}$ if 52.6 used | 2 |
| 9. | c |  | $\Delta U «=\frac{3}{2} n R \Delta T »=1.5 \times 243 \times 8.31 \times(19-(-12))=9.39 \times 10^{4}$ $Q «=\Delta U+W »=9.39 \times 10^{4}+6.27 \times 10^{4} \checkmark$ $Q=1.57 \times 10^{5} « \mathrm{~J} » \checkmark$ | Accept $1.60 \times 10^{5}$ if $66.1 \times 10^{3} \mathrm{~J}$ used Accept $1.55 \times 10^{5}$ if $61.6 \times 10^{3} \mathrm{~J}$ used | 3 |

(continued...)
(Question 9 continued)

| Question |  | Answers | Notal |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 9. | d | i | concave curve from RHS of present line to point above LHS of <br> present line $\checkmark$ <br> vertical line from previous curve to the beginning $\checkmark$ |  |
| 9. | d | ii |  |  |
| 9. |  |  |  |  |


| Question |  | Answers | Total |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 10. | a |  | $\frac{1}{2} \rho v_{X}^{2}=p_{Y}-p_{X}=\rho g \Delta h \checkmark$ <br> $v_{X}=\sqrt{2 \times 9.8 \times(0.32-0.10)} \checkmark$ <br> $v_{X}=2.08 \mu \mathrm{~ms}^{-1} » \checkmark$ | Notes |  |
| 10. | b | i | $R=« \frac{v r \rho}{\eta}=\frac{2.1 \times 0.25 \times 10^{3}}{8.9 \times 10^{-4}} » 5.9 \times 10^{5} \checkmark$ |  |  |
| 10. | b | ii | $(R>1000)$ flow is not laminar, so assumption is invalid $\checkmark$ | OWTTE |  |



## Option C - Imaging

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 12. | a | i | constructs ray parallel to principal axis and then to image position <br> OR <br> $u=8 \mathrm{~cm}$ and $\mathrm{v}=24 \mathrm{~cm}$ and lens formula | $e g:$ $2.0 \mathrm{~cm}$  <br> Allow answers in the range of 5.6 to 6.4 cm | 2 |
| 12. | a | ii | $m=$ «-»3.0 $\downarrow$ |  | 1 |
| 12. | b |  | completes diagram with blue focal point closer to lens $\checkmark$ blue light/rays refracted/deviated more <br> OR <br> speed of blue light is less than speed of red light $\checkmark$ <br> OR <br> different colors/wavelengths have different focal points/converge at different points $\checkmark$ | First marking point can be explained in words or seen on diagram | 2 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 13. | a | where the extensions of the reflected rays from the primary mirror would meet, with construction lines $\checkmark$ | eg: | 1 |
| 13. | b | greater magnification $\checkmark$ |  | 1 |
| 13. | c | «use of $\frac{1.22 \lambda}{d}$ to get» resolution of $6.7 \times 10^{-9}$ «rad» $\checkmark$ $\frac{5.8 \times 10^{-7}}{6.7 \times 10^{-9}}=87 « m » \downarrow$ <br> some reference to difficulty in making optical mirrors/lenses of this size $\checkmark$ | Allow $\frac{5.8 \times 10^{-7}}{5.5 \times 10^{-9}}=105$ «m » | 3 |

\begin{tabular}{|c|c|c|c|c|c|}
\hline \multicolumn{3}{|r|}{Question} \& Answers \& Notes \& Total <br>

\hline 14. \& a \& \& | $\sin c=\frac{1.4444}{1.4475} \text { or } \sin c=0.9978$ |
| :--- |
| critical angle $=86.2$ «`» $\checkmark$ |
| with cladding only rays travelling nearly parallel to fibre axis are transmitted |
| OR |
| pulse broadening/dispersion will be reduced | \& OWTTE \& 3 <br>

\hline 14. \& b \& i \& $$
\begin{aligned}
& \text { attenuation }=« 10 \log \frac{I}{I_{0}} »=10 \log \frac{2.0 \times 10^{-6}}{400 \times 10^{-6}} \\
& \text { attenuation }=«-» 23 « \mathrm{~dB} »
\end{aligned}
$$ \& Accept $10 \log \frac{400}{2.0}$ for first marking point \& 2 <br>

\hline 14. \& b \& ii \& | $185 \times 0.200=37$ loss over length of cable |
| :--- |
| « $\frac{37-23}{12}=1.17$ » so two amplifiers are sufficient | \& \& 2 <br>


\hline 14. \& b \& iii \& | mention of material dispersion mention that rays become separated in time |
| :--- |
| OR |
| mention that ray A travels slower/arrives later than ray B $\checkmark$ | \& \& 2 <br>

\hline
\end{tabular}

(continued...)
(Question 14 continued)

| Question |  | Answers | Notes |
| :---: | :--- | :--- | :--- | :--- |
| 14. | c | high bandwidth/data transfer rates $\checkmark$ <br> low distortion/Low noise/Faithful reproduction $\checkmark$ <br> high security $\checkmark$ <br> fast «fibre» broadband/internet $\checkmark$ <br> high quality optical audio $\checkmark$ <br> medical endoscopy $\checkmark$ | Allow any other verifiable sensible advantage |


| 15. | a | many/array of transducers send ultrasound through body/object $\checkmark$ <br> B scan made from many A scans in different directions $\checkmark$ <br> the reflection from organ boundaries gives rise to position $\checkmark$ <br> the amplitude/size gives brightness to the B scan $\checkmark$ <br> $2 D / 3 D$ image formed «by computer» $\checkmark$ | 3 max |
| :--- | :--- | :--- | :--- | :--- |

(continued...)
(Question 15 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 15. | b | i | the thickness of tissue that reduces the intensity «of the X -rays» by a half OR $x_{\frac{1}{2}}=\frac{\ln 2}{\mu}$ where $x_{\frac{1}{2}}$ is the half value thickness and $\mu$ is attenuation coefficient $\checkmark$ | Symbols must be defined for mark to be awarded | 1 |
| 15. | b | ii | after fat layer, $I_{\text {fat }}=I_{0} \mathrm{e}^{-0.4499 \times 5.00}$ <br> after muscle layer, $I=I_{\text {fat }} e^{-0.8490 \times 4.00} \checkmark$ $I=0.003533 I_{0} \text { or } 0.35 \% \checkmark$ |  | 3 |
| 15. | b | iii | «high energies factors:» <br> less attenuation/more penetration more damage to the body $\checkmark$ «so» stronger signal leaves the body <br> OR <br> «so» used in «most» medical imaging techniques $\checkmark$ «low energy factors:» <br> must be used with enhancement techniques $\checkmark$ greater attenuation/less penetration «so» more damage to the body «on surface layers» OR <br> «so» unwanted in «most» medical imaging techniques $\checkmark$ |  | 3 max |

## Option D - Astrophysics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 16. | a |  | photon/fusion/radiation force/pressure balances gravitational force/pressure <br> gives both directions correctly (outwards radiation, inwards gravity) $\checkmark$ | OWTTE | 2 |
| 16. | b |  | $« L \propto M^{3.5}$ for main sequence» luminosity of $P=2.5$ «luminosity of the Sun» $\checkmark$ |  | 1 |
| 16. | c | i | $\begin{aligned} & L_{\text {Gacrux }}=5.67 \times 10^{-8} \times 4 \pi \times\left(58.5 \times 10^{9}\right)^{2} \times 3600^{4} \\ & L_{\text {Gacrux }}=4.1 \times 10^{29} « \mathrm{~W} » \checkmark \\ & \frac{L_{\text {Gacrux }}}{L_{\odot}} «=\frac{4.1 \times 10^{29}}{3.85 \times 10^{26}} »=1.1 \times 10^{3} \end{aligned}$ |  | 3 |
| 16. | C | ii | if the star is too far then the parallax angle is too small to be measured OR <br> stellar parallax is limited to closer stars $\checkmark$ | OWTTE | 1 |

(continued...)
(Question 16 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 16. | d | i | line or area roughly inside shape shown - judge by eye $\checkmark$ | Accept straight line or straight area at roughly $45^{\circ}$ |  |
|  |  |  |  |  | 1 |
| 16. | d | ii | P between $1 L_{\odot}$ and $10^{1} L_{\odot}$ on main sequence drawn $\checkmark$ |  | 1 |

(continued...)
(Question 16 continued)

(continued...)
(Question 16 continued)

| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 16. | e | ALTERNATIVE 1 <br> Main sequence to red giant <br> planetary nebula with mass reduction/loss OR <br> planetary nebula with mention of remnant mass white dwarf $\checkmark$ <br> ALTERNATIVE 2 <br> Main sequence to red supergiant region $\checkmark$ <br> Supernova with mass reduction/loss <br> OR <br> Supernova with mention of remnant mass $\checkmark$ neutron star <br> OR <br> Black hole $\checkmark$ | OWTTE for both alternatives | 3 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 17. | a | use of gradient or any coordinate pair to find $H_{0} «=\frac{v}{d} »$ or $\frac{1}{H_{0}} «=\frac{d}{v} » \checkmark$ convert Mpc to m and km to m «for example $\frac{82 \times 10^{3}}{10^{6} \times 3.26 \times 9.46 \times 10^{15}}$ » $\checkmark$ age of universe «= $\frac{1}{H_{0}} »=3.8 \times 10^{17}$ «s» | Allow final answers between <br> $3.7 \times 10^{17}$ and $3.9 \times 10^{17}$ «S» or $4 \times 10^{17}$ «s» | 3 |
| 17. | b | non-accelerated/uniform rate of expansion OR $H_{0}$ constant over time $\checkmark$ | OWTTE | 1 |
| 17. | c | $\begin{aligned} & z «=\frac{v}{c} »=\frac{4.6 \times 10^{4} \times 10^{3}}{3.00 \times 10^{8}}=0.15 \\ & \frac{R}{R_{0}}=« z+1 »=1.15 \\ & \frac{R_{0}}{R}=« \frac{1}{1.15}=» 0.87 \end{aligned}$ <br> OR <br> $87 \%$ of the present size |  | 3 |


| Question |  | Answers | Notes |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 18. | a | $\begin{array}{l}\text { «For a star to form»: magnitude of PE of gas cloud }>\text { KE of gas cloud } \\ \text { OR } \\ \text { Mass of cloud }>\text { Jean's mass } \\ \text { OR } \\ \text { Jean's criterion is the critical mass } \checkmark \\ \text { hence a hot diffuse cloud could have KE which is too large/PE too small } \\ \text { OR } \\ \text { hence a cold dense cloud will have low KE/high PE } \\ \text { OR } \\ \text { a cold dense cloud is more likely to exceed Jeans mass } \\ \text { OR } \\ \text { a hot diffuse cloud is less likely to exceed the Jeans mass } \checkmark\end{array}$ | Accept $E_{p}+E_{k}<0$ |$]$| 2 |
| :--- |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 19. | a |  | «rotational» velocity of stars are expected to decrease as distance from centre of galaxy increases <br> the observed velocity of outer stars is constant/greater than predicted $\checkmark$ implying large mass on the edge «which is dark matter» $\checkmark$ | OWTTE <br> 1st and 2nd marking points can be awarded from an annotated sketch with similar shape as the one below | 3 |
| 19. | b |  | data from type 1a supernovae shows universe expanding at an accelerated rate gravity was expected to slow down the expansion of the universe <br> OR <br> this did not fit the hypotheses at that time $\checkmark$ <br> dark energy counteracts/opposes gravity <br> OR <br> dark energy causes the acceleration $\checkmark$ | OWTTE | 3 |

